



## Linking JPSS-1 and SNPP CrIS through Inter-calibration Efforts

#### Likun Wang<sup>\*1</sup>, Yong Chen<sup>1</sup>, Hui Xu<sup>1</sup>, Denis Tremblay<sup>2</sup>, Changyong Cao<sup>3</sup>

\* Univ. of Maryland, College Park, MD; <u>wlikun@umd.edu</u>
2. ERT, Inc., Laurel, MD
3. NOAA/NESDIS/STAR, College Park, MD











- JPSS and SNPP CrIS instrument characteristics
- Inter-comparison strategy
  - Direct comparison
  - In-direct comparison
- Conclusion Remarks



## **CrIS: Interferometer**







## **CrIS Spectral Bands**





## **CrIS Scan Patterns**





From Yong Han et al. (2014)

- For one mirror sweep, each focal planes illuminate 9 FOVs.
- The 9 FOVs form one FOR.
- Swath is 2200 Km (FOR1 to FOR 30).
- CrIS acquires 1 scan line every 8 seconds.
- CrIS measures 8.7 million spectra per day.



## J1 CrIS ICT Performance Greatly Improved From SNPP

- Additional PRT (two for SNPP) for J1 provided more temperature and gradient knowledge.
- Cavity wedge design increase ICT blackbody emissivity, eliminating views to other optical parts within instrument
- Simplify the ICT model and more accurate calibration performance is expected.



Figure 41: Radiometric model when the scene is the ICT

 $OMA(\sigma), T^{OMA}$ 



## **J1 NEdN Performance Equal to** or Better Than SNPP



#### J1 RRTVAC: Total NEdN ES Real NEdN (Radiances) Using Standard Deviation 1**0**⁰ FOV 1 Radiance (mW/m<sup>2</sup> sr cm<sup>-1</sup>) FOV 2 FOV 3 FOV 4 10<sup>-1</sup> FOV 5 FOV 6 FOV 7 FOV 8 FOV 9 10<sup>-2</sup> Spec 287K 10<sup>-3</sup>∟ 600

800 1000 1200 1400 1600 1800 2000 2200 2400 2600 Wavenumber (cm<sup>-1</sup>)

NPP TVAC4 MN: Total NFdN



- J1 LWIR and SWIR noise performance is similar to SNPP
- J1 MWIR FOV 9 is out of family but within specification
- SNPP MWIR FOV 7 is out of family but larger than specification



## J1 Detector Nonlinearity Levels comparable or better than SNPP

- J1 FOV 2 and 5 have largest nonlinearity in LWIR, compared to FOV 9 in SNPP
- J1 nonlinearity is greatly improved in MWIR, only FOV 9 shows nonlinearity.
- The nonlinearity will be further tuned on-orbit.



#### LWIR band









## <sup>cics</sup> Inter-calibration of SNPP and J1 CrIS

- Once J1 reaches in the final orbit, J1 will be the same orbit as SNPP except for the ½ orbit along track separation.
- SNO opportunities exist if instruments are turned on and collecting earth view data before orbit raising (Day 45).
- There will be NO SNOs between SNPP and J1 after it reaches final orbit.
- We need to explore the other ways to inter-calibrate J1 and SNPP CrIS.



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## Same orbits with 180° shift









From Denis JOUGLET

# cics Overlapped regions with time delay





- Similar viewing absolute angles (same atmospheric thickness)
- Opposite scan positions
- homogeneous and stable scenes in 50.7 min
  - Collocated VIIRS



## Satellite Zenith Angle for overlapped FOVs



Satelite Zenith Angle





## **Using IASI as a transfer radiometer**





It is very straightforward to convert CrIS radiances to overlapped GOES16 ABI channels.





OES16 ABI should be very stable within 50 minutes and very suitable to serve as a transfer radiometer.

A lot of samples for CrIS (SNPP and J1) and GOES16 ABI.

### cics-md **Using GPS RO as a transfer target**

BT (K)



the GPS RO retrieved profiles have shown potentials for use as a benchmark data set to validate other measurement at a range of 5–25 km





The previous study (Wang and Zou 2013) using GPS RO profiles to simulate SSU observations, showing very good agreement.





- Do we expect the BT difference along spectral domain between SNPP and J1 CrIS?
- If yes, are these BT differences trustable based on all direct and indirect comparison?
- If yes, do we understand the root causes of the BT differences?
- If yes, how can resolve these differences through reprocessing efforts to link SNPP and J1 for a decade of CrIS dataset?





# **QUESTIONS?**